# ESA multibody simulator for spacecrafts' ascent and landing in a microgravity environment 

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#### Abstract

The investigation of planets, moons and small bodies, including comets and asteroids can contribute substantially to our understanding of the formation and history of the solar system. In situ observations by landers play an important role in this field: for example, the Rosetta Lander Philae has been the first spacecraft to accomplish a soft touchdown on a comet. Since the last decade, the urgency to anticipate hardware performance and correlate testing results with mathematical models is generating the trend for more extensively applying multibody approach in supporting the design and verification of complex aerospace systems, from the early phases of the project. Some of the essential multibody technologies assisting the analysis of this class of problems consist of a reliable attitude control algorithm, a variable-step and variable-order numerical integrator and a consistent contact-friction formulation. This paper illustrates the particular implementation of these crucial features in the European Space Agency's multibody software DCAP and shows a direct application to a feasibility case study.


Keywords Multibody • DCAP • Landing • Touchdown • Quaternions • Integrator • Contact • Sloshing

## Abbreviations <br> CDF Concurrent design facility

[^0]| CoG | Centre of gravity |
| :--- | :--- |
| DCAP | Dynamic and control analysis package |
| DOF | Degree(s) of freedom |
| EOM | Equations of motion |
| ESA | European Space Agency |
| ESTEC | European Space Research and Technology |
|  | Centre |
| EXOMARS | Exobiology on Mars mission |
| GNC | Guidance, navigation and control |
| JAXA | Japan Aerospace Exploration Agency |
| JPL | Jet propulsion laboratory |
| NASA | National Aeronautics and Space |
|  | Administration |
| ODE | Ordinary differential equation |
| PECE | Prediction-evaluation-correction-evaluation |
|  | method |
| S/C | Spacecraft |
| TEC-MSM | ESA's mechanisms section |

## 1 Introduction

Exploring the diverse objects in our Solar System is the heart of space agencies scientific objectives, from the early manned exploration of the moon in the 60s to the most recent robotic missions to asteroids, comets and planets [1-3].

Few examples can effectively illustrate the diverse challenges of exploration missions: JAXA Hayabusa mission $[4,5]$ attempted to take samples from the surface of the asteroid Itokawa and successfully returned it to Earth in 2010; ESAs cornerstone mission Rosetta [6] has been the first to rendezvous with a comet and to deploy a lander (Philae) on the comets nucleus in 2014. In the last decade, NASA has successfully landed on Mars multiple rovers [7]
(Curiosity landed in 2012), ESA is adding to this list the EXOMARS missions [8] with Russian partnership and also China and India are planning exploration missions to the Earth's moon and the red planet before 2020. Moreover, the CleanSpace program [9] and other Earth low orbit debris removal missions such as e.Deorbit [10], currently in phase A, heavily rely on multibody simulations for both the complex close proximity capture operation and for analysis of the elastic dynamics during the de-orbiting burn.

Building on the 50 years heritage of the European Space Agency's Technology Center in providing effective spacecraft design and simulations technical solutions, this paper addresses the latest improvements implemented in the ESA multibody software DCAP, especially dedicated to ease simulating system dynamics during the landing of a spacecraft, in the frame of a feasibility study (CDF) and preliminary design verification activities. This paper builds on DCAP release 8.0, which includes extending the usability of the complementary features already discussed in recent papers by G. Baldesi et al. (e.g. launcher lift-off and multi-stage separation [11], integrated ascent and guidance approach [12]).

For the purpose of simulating a realistic descent and landing of a spacecraft, the complexity of three-dimensional motion has to be supported by a robust orientation algorithm capable of always describing the attitude of the body. In Sect. 2 of this paper, the Euler angles implemented in DCAP release 8.0 have been substituted with quaternions, to avoid numerical singularities for particular spacecraft orientations.

Touchdown technologies aim at providing effective solutions to unpredictable and/or occasional fast dynamics discontinuous events. Fixed-time step 4th order RungeKutta integrator, which is the one implemented in DCAP release 8.0 , is intrinsically limited in accurately simulating these type of events; as for other tools based on fixed-step integrators, the provision of an elaborate time reset and step-back subroutines logic has been able to partially compensate for the lack of automatic event detection and step adaptation in the DCAP package. Alternatively and more effectively, in the authors opinion, sudden discrete events can be comprehensively tackled by numerical integrators with variable time step and algorithm order, capable of effectively recognizing the discontinuity and automatically tune the solver parameters to match the overall required tolerance. Moreover, this family of integrators is known to exhibit outstanding computational speed. In the DCAP implementation proposed in Sect. 3, the additional advantage is to eliminate the need for the complicated time-reset and step-back interconnected logic management by directly delegating its function to the integrator itself. This enables to further increase the overall computational speed and accuracy, to automatically detect discrete events and to
simplify the tree of dependencies and management of the exceptions cases which, in return, also decreases the risk of incorrect outputs.

Adequate representation of the soil dynamic properties [13-15] as well as the lander structure [16] and shock absorbing characteristics [17] are key parameters for determining the equivalent friction between the lander and the soil. Also, a proper selection of a number of parameters [18] such as the vertical and lateral landing velocities and the spacecraft inclination angle of the landing target can determine the margin of risk to crash, bounce or tip over during landing. Considering the general objective of maintaining a synthetic and versatile approach, after reviewing more enhanced three-dimensional contact representations [19, 20] and their implementation in state-of-the-art commercial software packages, a trade-off between complex parameter management versus global effectiveness goes in favor of a simpler consolidated model. A straightforward contact formulation based on Hertz theory is therefore presented in Sect. 4.

An example of landing simulation is then presented in the Sect. 5. The adopted approach is typical of feasibility studies, such as those performed in the European Space Agency's Concurrent Design Facility (CDF), where a number of versatile conceptual models are used in parallel to check for design feasibility and requirements consistency, as well as to rapidly simulate and compare the foreseeable mission scenarios. Moreover, the comparison with the results obtained with an independent commercial software is also shown for completeness.

Finally, conclusions are drawn in the Sect. 6.

## 2 Quaternions

### 2.1 Background on quaternions

Euler angles are widely used in multibody software to describe the attitude of bodies by representing the orientation of a reference frame relative to another. Any orientation can be described by a sequence of three element rotations. There are 12 different sequence conventions of Euler angles.

In multibody simulations, Euler angles are also used during the numerical integration. The dynamic routine of a multibody software produces the accelerations of the system and by integrating them the entire kinematics of the system is obtained. Unfortunately, by referring only to rotational degrees of freedom for simplicity, it is possible to integrate the angular accelerations into angular velocities, but it is erroneous to integrate angular velocities to obtain the attitude values, because the angular velocity is not an exact differential [21]. Therefore, the equations of motion

Table 1 General pros and cons of Euler angles versus quaternions

| Euler angles | Quaternions |
| :--- | :--- |
| Easier to understand | Less computational overhead |
| Gimbal lock singularity | Difficult physical interpretation |
| Ambiguity in the angles sequence | Conversion to matrix more <br> efficient |

are generally formulated in terms of Euler angles and their derivatives, because it is allowed to double integrate the double dot Euler angles $\{\ddot{\theta}\}$ to obtain the Euler angles. The following relations convert Euler angles derivatives into angular velocities and accelerations:
$\{\omega\}=[L]\{\dot{\theta}\}$
$\{\dot{\omega}\}=[L]\{\ddot{\theta}\}+[\dot{L}]\{\dot{\theta}\}$
where the vectors $\{\omega\}$ and $\{\dot{\omega}\}$ contain the angular velocities and accelerations of a generic body, the vector $\{\dot{\theta}\}$ and $\{\ddot{\theta}\}$ contain the first and second Euler angles derivatives and the matrix $[L]$ and $[\dot{L}]$ are used for the conversion.

The main disadvantage of using the Euler angles is the numerical phenomenon of gimbal lock. In particular system configurations, the matrix [ $L$ ] becomes singular and the inverse relation can no more be extracted.

For example, assuming to use the Euler angles convention 1-2-3, the relation becomes:
$\{\omega\}=\left[L_{123}\right]\{\dot{\theta}\}$
where the matrix [ $L_{123}$ ] is defined as:
$\left[L_{123}\right]=\left[\begin{array}{lll}\cos \theta_{2} \cos \theta_{3} & \sin \theta_{3} & 0 \\ -\cos \theta_{2} \sin \theta_{3} & \cos \theta_{3} & 0 \\ \sin \theta_{2} & 0 & 1\end{array}\right]$
The inverse relation can been written as:
$\{\dot{\theta}\}=\left[L_{123}\right]^{-1}\{\omega\}$
where the matrix $\left[L_{123}\right]^{-1}$ is:
$\left[L_{123}\right]^{-1}=\left[\begin{array}{lll}\left(\frac{\cos \theta_{3}}{\cos \theta_{2}}\right) & -\left(\frac{\sin \theta_{3}}{\cos \theta_{2}}\right) & 0 \\ \sin \theta_{3} & \cos \theta_{3} & 0 \\ -\left(\frac{\cos \theta_{3} \sin \theta_{3}}{\cos \theta_{2}}\right) & \left(\frac{\sin \theta_{3} \sin \theta_{2}}{\cos \theta_{2}}\right) & 1\end{array}\right]$
For this particular Euler angles convention, the condition of gimbal lock occurs at Euler angle $\theta_{2}=90^{\circ}$, as the components of matrix $\left[L_{123}\right]^{-1}$ tends to infinity.

Introducing quaternions as an alternative to Euler angles is recognized as a valuable option to avoid the gimbal lock phenomenon. In this case, the relation of the angular velocity turns out to be [21]:
$\{\omega\}=2[Q]\{\dot{p}\}$
where the vector $\{\dot{p}\}$ is the first time derivative of the relative quaternions and the matrix $[Q]$ is defined as:

$$
=\left[\begin{array}{llll}
-p_{1} & p_{0} & p_{3} & -p_{2}  \tag{8}\\
-p_{2} & -p_{3} & p_{0} & p_{1} \\
-p_{3} & p_{2} & -p_{1} & p_{0}
\end{array}\right]
$$

and $p_{i}$ are the four quaternions defined as:

$$
\begin{align*}
& p_{0}=\cos \frac{\theta}{2} \quad p_{1}=u_{x} \sin \frac{\theta}{2} \quad p_{2}=u_{y} \sin \frac{\theta}{2} \\
& p_{3}=u_{z} \sin \frac{\theta}{2} \tag{9}
\end{align*}
$$

where $\theta$ is the rotation about the screw axis $\hat{u}$.

### 2.2 Quaternions implementation in DCAP

The $\operatorname{Order}(\mathrm{n})$ algorithm [11], used by DCAP release 8.0 to evaluate the dynamics of the system, produces equations of motion in terms of Euler angles derivatives. The algorithm treats separately all the hinges of the system, which are the connections between bodies. In this framework, only hinges with 3 rotational degrees of freedom could potentially experience gimbal lock.

The quaternions implementation has been designed so that, although DCAP will rely on quaternions for computing the kinematics of the system, Euler angles will still always be produced as standard output angles, because of their direct physical interpretation. This architecture achieves the known benefits from a computational prospective, see Table 1, while hiding the only main drawback which is the lack of intuitiveness from a user prospective.

The recoded internal symbolic generator has been enabled to assemble the equations of motion using the angular velocities $\{\omega\}$ and accelerations $\{\dot{\omega}\}$, instead of $\{\dot{\theta}\}$ and $\{\ddot{\theta}\}$, as generalized coordinates.

Considering an open loop topology system, it is usual to refer to the body $L(j)$ as the lower body of the body $j$, within the topology tree. Using Euler angles the following equations are used in DCAP:
$\left\{{ }^{L(j)} \omega^{j}\right\}=\left[L^{j}\right]\left\{\dot{\theta}^{j}\right\}$
$\left\{{ }^{L(j)} \dot{\omega}^{j}\right\}=\left[L^{j}\right]\left\{\ddot{\theta}^{j}\right\}+\left\{{ }^{L(j)} \stackrel{\circ}{\omega}^{j}{ }^{j}\right\}$
where $\left\{{ }^{L(j)} \omega^{j}\right\}$ is the relative angular velocities vector, $\left\{{ }^{L(j)} \dot{\omega}^{j}\right\}$ is the relative angular accelerations vector and $\left\{{ }^{L(j)}{ }_{\omega}^{\omega_{v}}{ }^{j}\right\}$ is the remainder terms vector between the body $j$ and its lower body $L(j)$.

From this, it is possible to compute the equations of motion in the same way and using the same routine, but considering physical entities such as angular velocities $\left\{{ }^{L(j)} \omega^{j}\right\}$ and angular accelerations $\left\{{ }^{L(j)} \dot{\omega}^{j}\right\}$, instead of the Euler angles derivatives, by means of the following substitution:
$\left[L^{j}\right]=[I]$
$\left\{L(j) \stackrel{\circ}{\omega}_{v}{ }^{j}\right\}=\{0\}$
where [ $I$ ] is the identity matrix. In this way, the vectors $\left\{{ }^{L(j)} \omega^{j}\right\}$ and $\left\{{ }^{L(j)} \dot{\omega}^{j}\right\}$ are equal to the vectors $\left\{\dot{\theta}^{j}\right\}$ and $\left\{\ddot{\theta}^{j}\right\}$ and they can be used as generalized coordinates. This simple mathematical substitution allows the rest of the DCAP core algorithm to be unchanged and avoids the complete reconditioning of the equations of motion generator algorithm.

Considering only rotational DOF for simplicity, since the EOM are derived using Kane's method of generalized speeds [22], the partial velocity matrices can be written as:
$\left[V^{j}\right]=\frac{\partial\left\{\dot{R}^{j}\right\}}{\partial\left\{{ }^{L(j)} \omega^{j}\right\}}$
$\left[\omega^{j}\right]=\frac{\partial\left\{\dot{\omega}^{j}\right\}}{\partial\left\{{ }^{L(j)} \omega^{j}\right\}}$
where $\left\{\dot{R}^{j}\right\}$ is the linear velocity vector of the CoG of the body $j$ and $\left\{\dot{\omega}^{j}\right\}$ is the angular velocity vector of the body $j$.

These partial velocity matrices will be automatically computed using the aforementioned substitutions (12) and (13) and no additional modifications to the EOM generator are necessary.

The state vector derivative between two bodies $L(j)$ and $j$, in the original DCAP algorithm, is:
$\left\{{ }^{L(j)} \dot{X}^{j}\right\}=\left\{\begin{array}{l}\left\{{ }^{L(j)} \ddot{\theta}^{j}\right\} \\ \left\{^{L(j)} \dot{\theta}^{j}\right\}\end{array}\right\}$
while, after the aforementioned substitution, it can be expressed as follows:
$\left\{^{L(j)} \dot{X}^{j}\right\}=\left\{\begin{array}{l}\left\{{ }^{L(j)} \dot{\omega}^{j}\right\} \\ \left\{^{L(j)} \omega^{j}\right\}\end{array}\right\}$
Before integrating, the state vector derivative is then transformed into:

$$
\left\{{ }^{L(j)} \dot{X}^{j}\right\}=\left\{\begin{array}{c}
\left\{{ }^{L(j)} \dot{\omega}^{j}\right\}  \tag{18}\\
0.5\left[^{L(j)} Q^{i}\right]^{T}\left\{{ }^{L(j)} \omega^{j}\right\}
\end{array}\right\}=\left\{\begin{array}{l}
\left\{^{L(j)} \dot{\omega}^{j}\right\} \\
\left\{^{L(j)} \dot{p}^{j}\right\}
\end{array}\right\}
$$

where:

- the vector $\left\{^{L(j)} \dot{p}^{j}\right\}$ contains the derivatives of the quaternions:

$$
\left\{{ }^{L(j)} \dot{p}^{j}\right\}^{T}=\left\{\begin{array}{llll}
L(j) & \dot{p}_{0}^{j} & L(j) \dot{p}_{1}^{j} & L(j) \\
\dot{p}_{2}^{j} & L(j) & \dot{p}_{3}^{j}
\end{array}\right\}
$$

- the matrix $\left[{ }^{L(j)} Q^{j}\right]$ is used to convert quaternions into angular velocity:

$$
\left[{ }^{L(j)} Q^{j}\right]=\left[\begin{array}{llll}
-^{L(j)} p_{1}{ }^{j} & { }^{L(j)} p_{0}{ }^{j} & { }^{L(j)} p_{3}{ }^{j} & -^{L(j)} p_{2}{ }^{j} \\
-^{L(j)} p_{2}{ }^{j} & -^{L(j)} p_{3}{ }^{j} & { }^{L(j)} p_{0}{ }^{j} & { }^{L(j)} p_{1}{ }^{j} \\
-^{L(j)} p_{3}{ }^{j} & L^{(j)} p_{2}{ }^{j} & -^{L(j)} p_{1}{ }^{j} & L^{L(j)} p_{0}{ }^{j}
\end{array}\right] .
$$

After integrating, the state vector becomes:
$\left\{{ }^{L(j)} X^{j}\right\}=\left\{\begin{array}{l}\left\{{ }^{L(j)} \omega^{j}\right\} \\ \left\{^{L(j)} p^{j}\right\}\end{array}\right\}$
Each quaternion must be normalized by evaluating the norm of the vector and dividing each component by the norm:

$$
\begin{gather*}
{ }^{L(j)}{\hat{p_{0}}}^{j}=\frac{{ }^{L(j)} p_{0}^{j}}{\left|\left\{{ }^{L(j)} p^{j}\right\}\right|}, \quad L(j){\hat{p_{1}}}^{j}=\frac{{ }^{L(j)} p_{1}^{j}}{\left|\left\{{ }^{L(j)} p^{j}\right\}\right|} \\
L^{L(j)} \hat{p}_{2}^{j}=\frac{{ }^{L(j)} p_{2}^{j}}{\left|\left\{{ }^{L(j)} p^{j}\right\}\right|}, \quad L(j) \hat{p}^{j}=\frac{{ }^{L(j)} p_{3}^{j}}{\left|\left\{{ }^{L(j)} p^{j}\right\}\right|} \tag{20}
\end{gather*}
$$

where ${ }^{L(j)} \hat{p}_{i}^{j}$ is the normalized quaternion.
As the non-linear time simulation algorithm requires Euler angles to evaluate the transformation matrices between bodies, also the modified algorithm retrieves the Euler angles at each time step from the quaternions [23].

### 2.3 Example of gimbal lock and quaternions results

Figure 1 shows a simple system composed by one body linked with a spherical hinge ( 3 rotational DOF) to the ground. The body 1 is initially rotated of $85^{\circ}$ around the second Euler angle axis and an initial velocity $\omega$ of $10 \%$ s is imposed. The Euler angles sequence 1-2-3 is chosen for this example, which exhibits gimbal lock at $90^{\circ}$.

Two simulations are compared: one using the original DCAP with Euler angles and one using the modified DCAP with quaternions.

The results are reported in Fig. 2: as expected, when the body 1 reaches the gimbal lock configuration $\left(\theta_{2}=90^{\circ}\right)$ the numerical integration based on Euler angles crashes, while the one based on quaternions seamlessly progresses through.


Fig. 1 Topology of the system for the gimbal lock example

Fig. 2 The evolution of the second Euler angle in both Euler angles case and quaternions case is reported on the left, the quaternions evolution is shown on the right, for the modified DCAP simulation


## 3 PECE integrator

The ODE solver, selected for DCAP, was developed by Shampine and Gordon [24] in collaboration with Sandia Laboratories and the University of New Mexico in 1974. It is based on a PECE Adams-BashforthMoulton algorithm with variable-step size and varia-ble-order algorithm, and the code is written in Fortran language.

### 3.1 Background on PECE algorithm

A PECE algorithm is a Prediction-Evaluation-Correc-tion-Evaluation method and both explicit and implicit methods are used. The sequence of actions are here summarized:

- Prediction an explicit Adams-Bashforth method of order $k$ is used to predict the value $y_{n+1}$;
- Evaluation the function $f_{n+1}^{p}$ at $t_{n+1}$ is evaluated by means of the value $y_{n+1}$;
- Correction an implicit Adams-Moulton method of order $k+1$ is then used to correct the previous value and to obtain a more accurate estimation of the solution, $y_{n+1}$;
- Evaluation another evaluation of the function $f_{n+1}$ is carried out.

Adams method approximates the derivative with a polynomial interpolating the computed derivative values and then integrates the polynomial. The variable step size is able to follow the slow or fast modifications of the function during the integration.

The Adams-Bashforth method for variable step size is:
$p_{n+1}=y_{n}+h_{n+1} \sum_{i=1}^{k} g_{i, 1} \phi_{i}^{*}(n)$
where $p_{n+1}$ stands for the predictor of the PECE algorithm and $k$ is the order of the method. Quantities in Eq. (21) are defined as [24]:
$\phi_{i}^{*}(n)=\beta_{i}(n+1) \phi_{i}(n)$
$\phi_{1}(n)=f\left[x_{n}\right]$
$\phi_{i}(n)=f\left[x_{n}, x_{n-1}, \ldots, x_{n-i+1}\right] \prod_{j=2}^{i} \Psi_{j-1}(n) \quad i>1$
$\Psi_{i}(n+1)=\sum_{j=1}^{i} h_{n+2-j}$
where $h_{i}$ is the time step and $\alpha_{i}(n)$ and $\beta_{i}(n)$ are defined as:
$\alpha_{i}(n+1)=\frac{h_{n+1}}{\Psi_{i}(n+1)} \quad i \geq 1$
$\beta_{1}(n+1)=1$
$\beta_{i}(n+1)=\frac{\prod_{j=2}^{i} \Psi_{j-1}(n+1)}{\prod_{j=2}^{i} \Psi_{j-1}(n)} \quad i>1$
and the coefficient $g_{i, 1}$ is identified by:
$g_{1, q}=\frac{1}{q}$

Table 2 General pros and cons of PECE integrator compared to fixed Runge-Kutta integrator

| Runge-Kutta 4th | PECE |
| :--- | :--- |
| Deterministic simulations | Faster |
| Easier implementation | More stable and efficient |
| Easier to use in a control loop | Greater order of the algorithm |
| Requires very little internal data storage per iteration | Possibility of finding an estimate of the error at each time step |
|  | Detection of sudden discontinuities not known a priori |

$g_{2, q}=\frac{1}{q(q+1)}$
$g_{i, q}=g_{i-1, q}-\alpha_{i-1}(n+1) g_{i-1, q+1} \quad i \geq 3$
where the parameter is an integer number $q=1,2,3 \ldots$
The Adams-Moulton method with variable step size uses the same data of the predictor method, plus an additional one, accounting for the order $k+1$ of the corrector, therefore:
$y_{n+1}=p_{n+1}+h_{n+1} g_{k+1,1} \phi_{k+1}^{p}(n+1)$
The main advantage of this formula is the possibility to evaluate $g_{k+1,1}$ during the prediction process. In this case, the last factor in (32) is defined as:
$\phi_{i+1}^{p}(n+1)=\phi_{i}^{p}(n+1)-\phi_{i}^{*}(n)$
Since the algorithm of order $k$ is not self-starting, it is always necessary to start from a first order method and than increase the order up to the desired one.

The ODE algorithm identifies the adequate step size able to minimize the local error, which translates into approximating the local solution uniformly well over the entire interval. The local error estimation procedure is accomplished by comparing the results of formulas (21) and (32) at different orders [25].

By considering a predictor of order $k$ and a corrector of order $k+1$
$y_{n+1}(k)=p_{n+1}+h_{n+1} g_{k, 1} \phi_{k+1}^{p}(n+1)$
$y_{n+1}(k+1)=p_{n+1}+h_{n+1} g_{k+1,1} \phi_{k+1}^{p}(n+1)$
the local error estimator becomes:
$l e_{n+1}(k)=h_{n+1}\left(g_{k+1,1}-g_{k, 1}\right) \phi_{k+1}^{p}(n+1)$
where $l e_{n+1}(k)$ stands for the local error at the step $n+1$ with an order $k$ algorithm.

In the same manner, the local error is then estimated by considering various orders: for a simpler and easier discrimination of the optimal order, the step size is assumed to be constant during all the local error evaluations.

To limit the mesh distortion, the step size $h_{n+1}$ is imposed to vary only between $\frac{1}{2} h_{n}$ and $2 h_{n}$, therefore a


Fig. 3 System composed of a pendulum and a surface
constant step size $h$ is a reasonable choice to represent the interval. For the same reason, ODE algorithm is not suitable for managing differential equations where drastic step size reductions are necessary (stiff problem).

### 3.2 PECE implementation in DCAP

The ODE subroutines have been coded inside the DCAP source in a parallel stream, so that the user has the option to chose from two independent solvers: the traditional fixed-step Runge-Kutta 4th order integrator and the new ODE one. Table 2 reports the main pros and cons between a Runge-Kutta fixed step size and 4th order integrator and a PECE integrator.

As part of the ODE implementation in the DCAP core, bypassing the DCAP release 8.0 time reset and step-back features has proven mandatory throughout the code, to enable the ODE algorithm itself to directly detect sudden discontinuities and consequently adapt the step size and algorithm order. Conversely, the time reset and stepback features remain fully interlinked to the Runge-Kutta stream.

### 3.3 Example of ODE results

The topology of a simple system composed by one body is displayed in Fig. 3. The body is linked to the ground by means of a 1 DOF rotational joint and a damped spring in the same direction. The possibility of elastic contact (topic

Fig. 4 Angle of a pendulum which hits a surface


Fig. 5 Time step and order $k$ of the PECE algorithm of the ODE solver for the first 10 s of the simulation

of next chapter) is foreseen between the $P$ point of the body and the surface $S$. The body initially is aligned with respect to the $y$ axis and an initial angular velocity of $50 \%$ s is imposed.

Figure 4 reports the results of two simulations: one uses the Runge-Kutta 4th order solver with a fixed time step of $10^{-4} \mathrm{~s}$ and the other relies on the ODE code with a tolerance of $10^{-6}$. As expected, the ODE solver detects the event of the contact and reduces the time step and the order of the PECE algorithm, as demonstrated also by Fig. 5.

A comparison of the computational time speed, executing DCAP on the same machine, shows that the PECE solver is $38 \%$ faster than the Runge-Kutta one, in this specific simulation case.

## 4 Contact

In the aerospace field, offering a range of contact modelling is of crucial relevance for multibody software which aim at tackling the landing of spacecraft on unknown gravel surfaces. Unfortunately, DCAP release 8.0 was lacking this capability, as it only incorporates a basic elastic contact algorithm between one point and an infinite plane, without soil damping properties and without any tangential friction implementation.

### 4.1 Contact implementation in DCAP

With the objective of better assisting preliminary assessment activities, priority has been given to enhancing


Fig. 6 Contact type "point to sphere"
and developing the following three fundamental contact configurations:

- "point to infinite plane" contact;
- "point to finite plane" contact;
- "point to sphere" contact.

All of them implement the same Hertzian contact formulation, based on the following equations:

- for the normal direction:
$F_{n}=k \cdot \delta^{3 / 2}+c \cdot \frac{\mathrm{~d} \delta}{\mathrm{~d} t}$
- for the tangent directions:
$F_{t}=F_{n} \cdot \mu$
where $F_{n}$ and $F_{t}$ are the normal and tangent forces, $k$ is the stiffness, $\delta$ is the penetration depth, $c$ is the damping and $\mu$ is the friction coefficient.

Figure 6 displays that the contact vector is always pointing from the reference node on the surface (called contact point) to the reference node of the body (called tip): for planar contact types, the vector is perpendicular to the plane, while the vector is normal to the surface of the sphere for the spherical contact.

Figure 5 displays how, in the case of ODE, the detection of the contact is automatic and the computational accuracy depends on the integrator tolerance: once the algorithm


Fig. 7 Lander configuration

Table 3 Pads positions with respect to the $\mathrm{S} / \mathrm{C} \mathrm{CoG}$

| - | $\mathrm{x}(\mathrm{m})$ | $\mathrm{y}(\mathrm{m})$ | $\mathrm{z}(\mathrm{m})$ |
| :--- | ---: | :--- | ---: |
| Pad 1 | 0.8877 | -1.2968 | 1.4935 |
| Pad 2 | 0.8877 | -1.2968 | -1.4935 |
| Pad 3 | -1.7001 | -1.2963 | 0.0000 |

recognizes the contact event, the integration time step and the order of the algorithm are adapted until the tolerance is matched.

## 5 Landing simulation

In this Section, the case study of a spacecraft landing on the surface of an asteroid is presented, where the DCAP model is built by taking advantage of all the additional capabilities presented in this paper.

To benchmark and validate the simulation results, an equivalent model has been also built into an independent commercial software (e.g., MSC Adams), where the system has the same properties, and a "point to plane" contact is adopted for the impact with the surface for simplicity.

The spacecraft, as shown in Fig. 7, is composed of a total of five rigid bodies connected with each other: a S/C main body, one sloshing pendulum and three legs. The main body of the spacecraft is linked to the legs by means of 1 DOF translational joints embedding crushable devices. On top of the spacecraft, four thrusters are mounted to push down the lander, after the impact with the surface, with a force of $5 N$ per thruster.

The total mass of the lander is 362 kg , and the inertia matrix, with respect to its inertial reference frame, is:
$[I]=\left[\begin{array}{lll}645.8 & 6.2 & -14.9 \\ 6.2 & 690.1 & 9.1 \\ -14.9 & 9.1 & 628.6\end{array}\right] \mathrm{kg} \mathrm{m}^{2}$

Fig. 8 Pads positions with respect to the spacecraft CoG


Table 3 gives the location of the pads with respect to the CoG of the S/C and Fig. 8 displays two images of the S/C with respect to two perpendicular planes; the overall footprint diameter is about 3.4 m .

Crushable devices are designed to reduce the shock transmitted to the upper part of the spacecraft by absorbing the impact with the terrain [17, 26, 27] and they are modeled as a Coulomb damper, which globally represent
its elasto-plastic deformation. The crushable positions and orientations with respect to the spacecraft CoG are reported in Table 4 and shown in Fig. 8, along the legs of the lander.

Figure 9 shows the characteristics of the crushable devices in terms of displacements and forces. Each device is designed to apply a constant force in the range of 400 N at the interface of the leg to the spacecraft, with a maximum displacement limited to about 20 cm .

Table 4 Crushable reference frames positions and orientations with respect to the S/C CoG

| - | $\mathrm{x}(\mathrm{m})$ | $\mathrm{y}(\mathrm{m})$ | $\mathrm{z}(\mathrm{m})$ | $\theta_{1}\left({ }^{\circ}\right)$ | $\theta_{2}\left({ }^{\circ}\right)$ | $\theta_{3}\left({ }^{\circ}\right)$ |
| :--- | :--- | :--- | :--- | ---: | :--- | ---: |
| Crushable 1 | 0.8668 | -1.0000 | 1.4550 | -7.4 | 0.0 | 4.0 |
| Crushable 2 | 0.8668 | -1.0000 | -1.4550 | 7.4 | 0.0 | 4.0 |
| Crushable 3 | -1.6532 | -1.0000 | 0.0000 | 0.0 | 0.0 | -9.0 |



Fig. 9 Crushable displacements and forces


Fig. 10 Predicted linear and angular acceleration of the spacecraft

The sloshing effect implemented in this DCAP model aims at introducing an equivalent dynamic disturbance to the spacecraft motion at impact, and does not pretend to accurately encompass all the complex fluid-dynamics of sloshing phenomena. Therefore, following simple representations from literature [28], an equivalent pendulum mass is linked to the $\mathrm{S} / \mathrm{C}$ CoG by means of a torsional springs. The remaining fuel, contributing to the sloshing effect, is estimated 10 kg . The position of the pendulum, as seen in Fig. 8, is at 26.8 mm on $Y$ axis with respect to the S/C CoG with a pendulum length of 150 mm ; stiffness and damping values are 10 and $1 \mathrm{~N} / \mathrm{m}$.

A contact model "point to sphere" between the pad of each leg and the surface of the asteroid represents a realistic model of the touchdown. Therefore, the pad has been modeled as a simple point and the CAD pad representation is only for visual effect. Although a fine characterisation of the soil is crucial to obtain more realistic simulations, the lack of data in early design phases lead to use generic values from internal knowhow.

Moreover, the DCAP gravity gradient feature [11] has also been used to include the weak gravitational field of the asteroid body.

At touchdown, the spacecraft is initially rotated of about $10^{\circ}$ with respect to the surface and the S/C CoG is at about 2.4 m from it. The leg number 3, as shown in Fig. 7, is the first one to get in contact with the surface.

Figure 10 shows the linear and angular accelerations of the spacecraft during the landing phase. The maximum linear acceleration peak is less than $2.5 \mathrm{~m} / \mathrm{s}^{2}$, which is found in line with requirements in similar cases [29].


Figure 11 compares the linear and angular velocities of the spacecraft evaluated using DCAP and MSC Adams. Both simulations are found in good agreement: at touchdown, the peak values reach about $0.96 \mathrm{~m} / \mathrm{s}$ for the linear velocity, and $0.43 \mathrm{~m} / \mathrm{s}$ for the angular velocity, which is found acceptable.

The resulting forces acting on each leg of the S/C are compared in Fig. 12 for both DCAP and MSC Adams models. The overall dynamics of the system during touchdown is well reproduced. The minor differences are due to the contact type assumed for the two multibody models, as the evolution of the local relative impact angle of each pad is slightly different in case of spherical or planar surface. The sloshing mass, which keeps moving for few seconds after the landing is completed, is the main responsible for the oscillations visible in each graph.

## 6 Conclusion

In conclusion, this paper initially summarizes the most critical features essential for touchdown and landing simulations, and focuses on their particular implementation and validation experience inside the European Space Agency's DCAP software.

Although quaternions avoid the risk of gimbal lock and prove computationally more efficient than Euler angles, the latter ones are still conveniently displayed to users as the standard output in DCAP, thanks to their direct link to physics.

With respect to fixed-step integrators, ODE solvers based on PECE algorithm provide users with faster


Fig. 11 Linear and angular velocities of the spacecraft in DCAP and MSC Adams


Fig. 12 Contact forces during the touchdown of the spacecraft for both DCAP and Adams models
solutions and improved accuracy and allow the software developer to harmonize the internal processes by simplify the management of sudden events.

Also, the paper shows that simple contact algorithms combined into a consistent space multibody simulator can adequately provide technical support to the preliminary design activities, such as feasibility studies in a concurrent environment.

As further evidence, the performance of a DCAP-based simulator, including the features introduced in this paper, is compared to MSC Adams in a realistic landing simulation: the outcome confirms that the multibody simulator based on the DCAP tool is computationally advantageous and technically adequate for this type of activities.

Finally, this paper intends to keep raising awareness in the space community towards the advantages of introducing
the multibody approach as a fundamental tool for planetary exploration, whenever complex dynamics is involved.

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